

CERES Programmable Azimuth for Instrument Inter-calibration and Coverage of Field Campaigns

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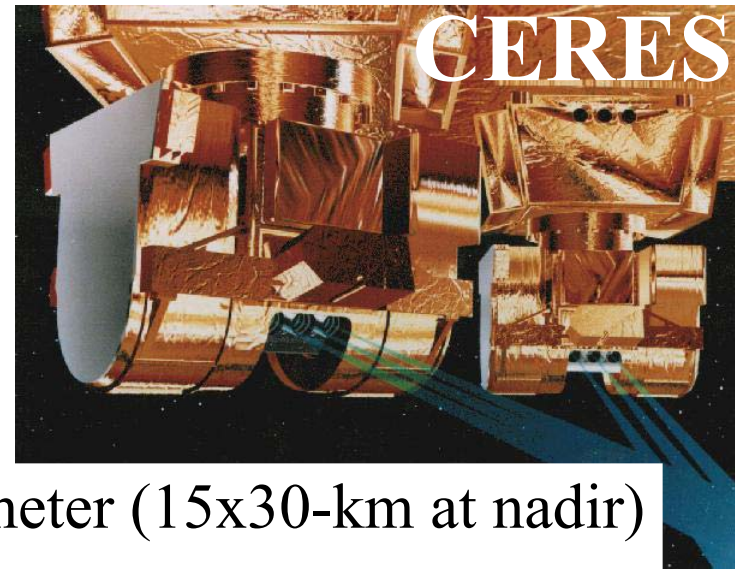
Science Applications International Corporation, VA/US

CERES/GERB Science Team Meeting

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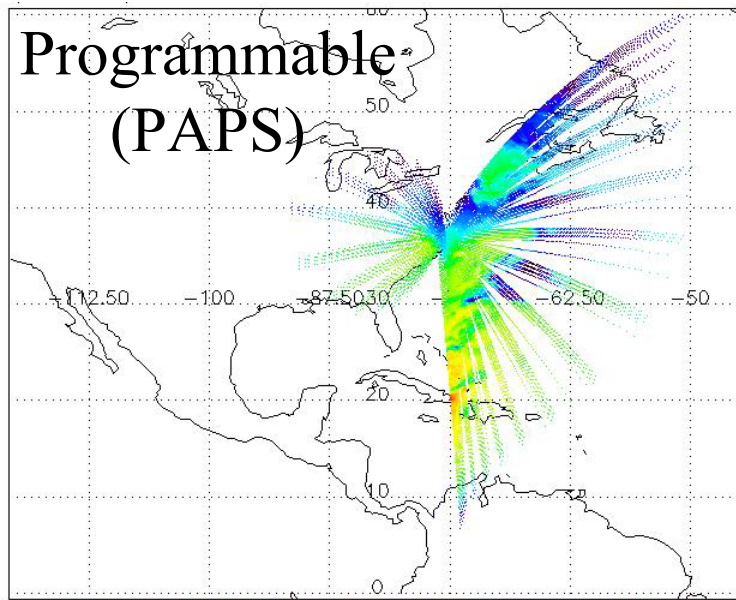
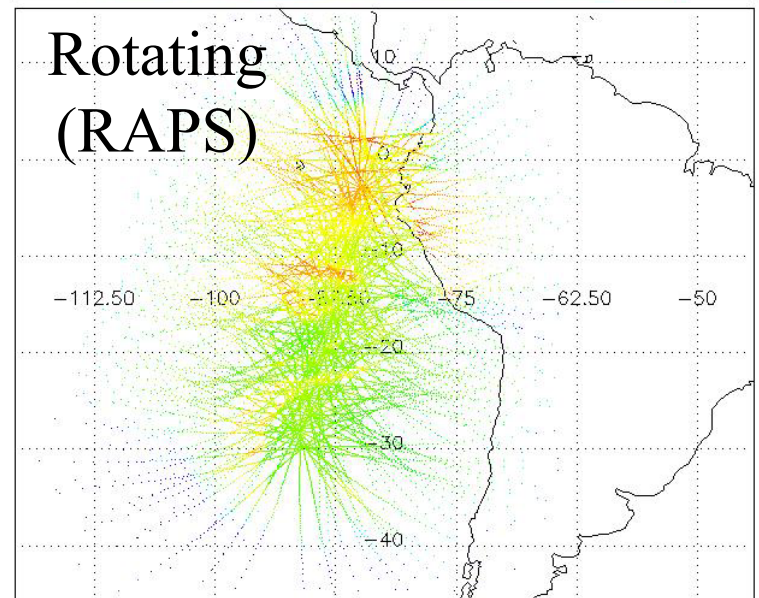
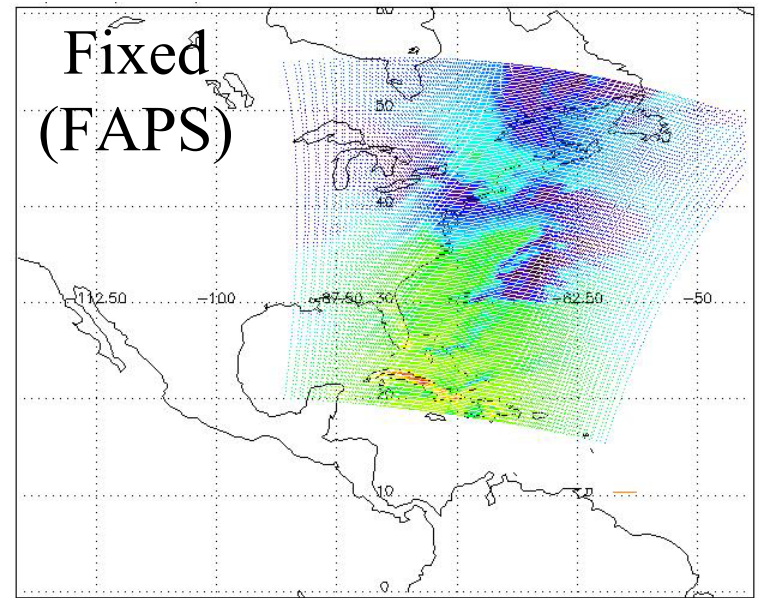
Cloud and the Earth's Radiant Energy System



- Narrow field-of-view scanning radiometer (15x30-km at nadir)
- Measures radiances in 0.2-4 μ m 0.2-100 μ m and 8-12 μ m
- Calibration stability monitored with
 - On-board calibration sources (blackbodies, lamps, solar)
 - Multi-channel and multi-instrument consistency
 - Geophysical calibration
- Gain drifts can be detected at the 0.1% level
- CERES/Terra: 0.25% LW, 0.1% SW and 0.1% WW (per year)

CERES

Azimuth Plane Scan Modes



1.90 2.67 3.44 4.21 4.98 5.75 6.52 7.29 8.06 8.83 9.60
Watts per square meter per steradian per micron

0.60 1.15 1.70 2.25 2.80 3.35 3.90 4.45 5.00 5.55 6.10
Watts per square meter per steradian per micron



PAPS for Inter-Calibration

CERES scanning plane is programmed to match the viewing geometry of other instruments:

- Instrument inter-calibration
- Direct comparisons of measured radiances
- Difference due to error in gain and unfiltering process
- Comparison uncertainty dominated by spatial noise

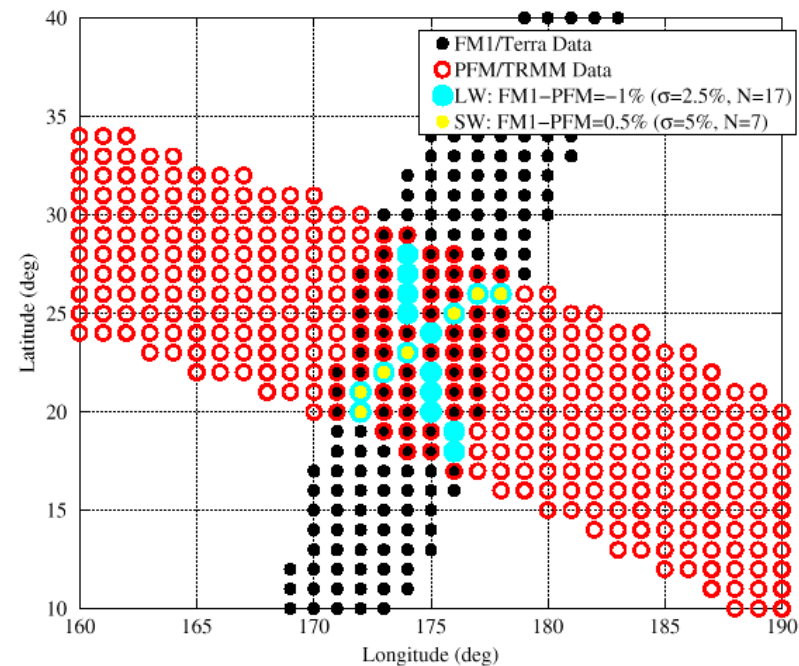
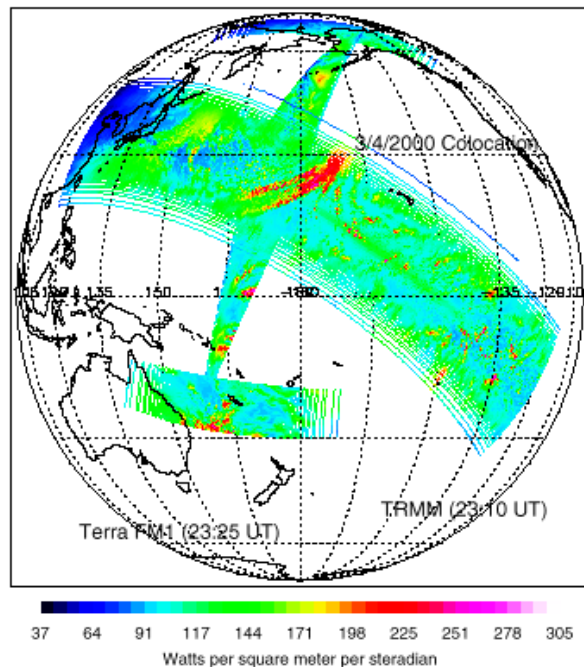
Examples:

- CERES/TRMM vs ScaRaB/Resurs
- CERES/TRMM vs CERES/Terra



LEO Satellite Inter-Calibration

- Satellite orbital periods must be different (sun-sync & precessing)
- Scan planes aligned at orbital crossing to match azimuth angle
- Collocated data with matched viewing zenith angle
- Main source of uncertainty is spatial noise



Previous LEO Satellite Inter-Comparisons

- Each orbital crossing is an independent sample
- Uncertainty $\mathcal{E} = \frac{t_{\alpha/2}\sigma}{\sqrt{N}}$
- Spatial noise dominates
- 100 independent samples
- $\sigma(\text{SW}) \times 4\sigma(\text{LW})$

(Wm ⁻² sr ⁻¹)		Δ	σ	N	\mathcal{E}
CERES-ScaRaB	SW	1.1	2.2	26	0.9
	LW	-0.5	0.5	50	0.2
TRMM-Terra	SW	-0.3	2.2	120	0.4
	LW	0.1	0.4	120	0.1

PAPS for Angular Model Validation

1. CERES scanning plane programmed to provide multi-angle observations of a pre-defined target:
 - Multiply spatial coverage of target area by 10
 - Reduce uncertainty of instantaneous radiant fluxes
 - Compare measured and modeled BDRF
2. Intensive observations of particular azimuth planes
 - Study anisotropy in principal plane (clear ocean)



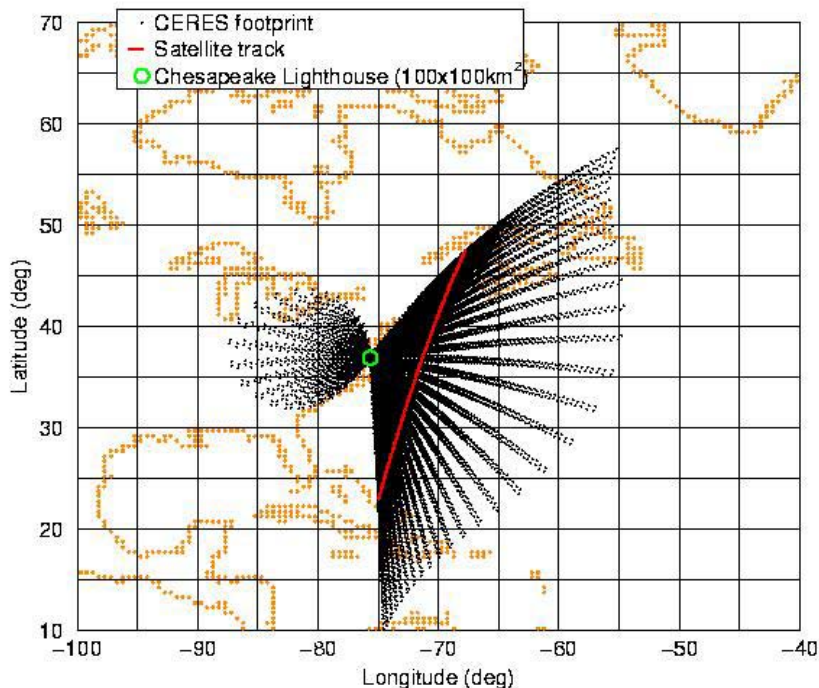
Method

- Define target location, or azimuth plane, or viewing geometry to match
- Get orbital prediction of sub-satellite location for dates of interest
- Compute CERES azimuth angles required to provide desired viewing geometry
- Produce command file for CERES upload

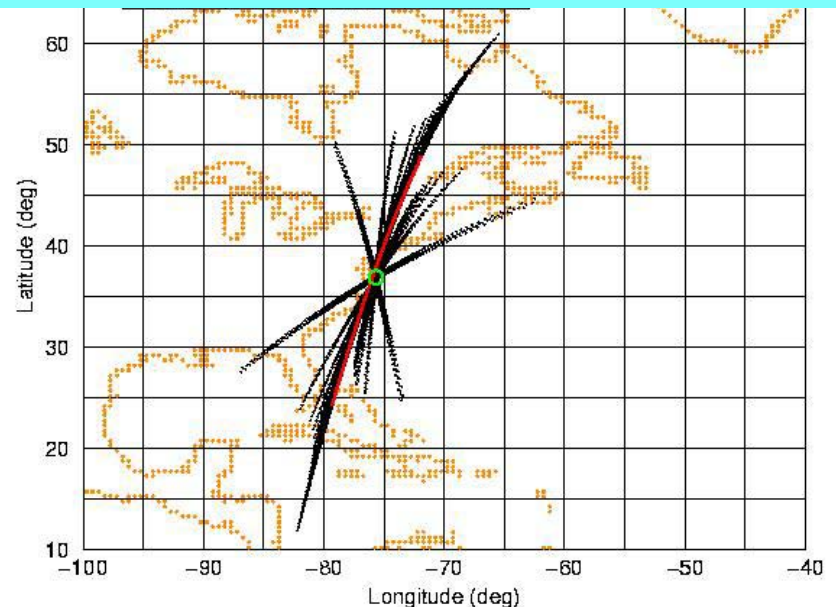
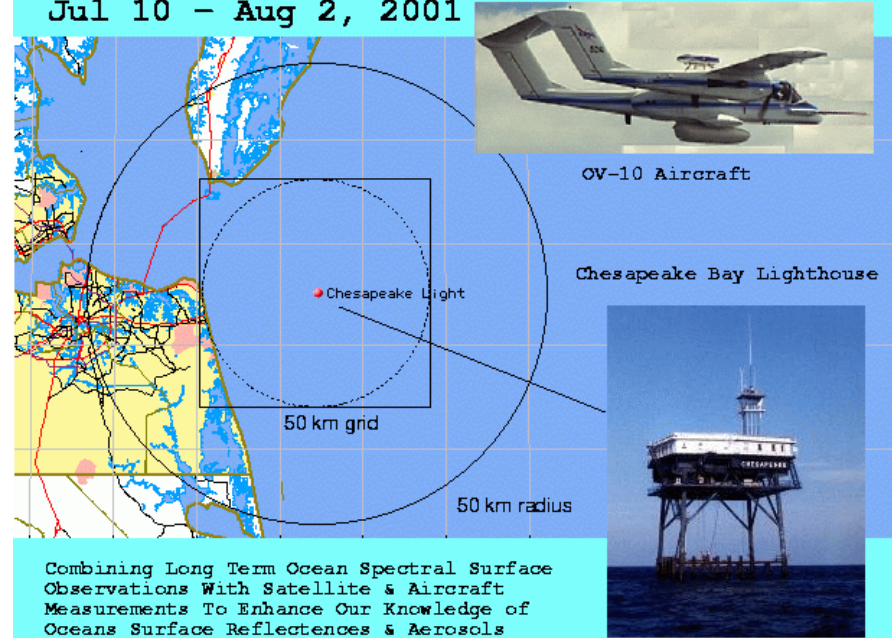


CERES Operations during CLAMS

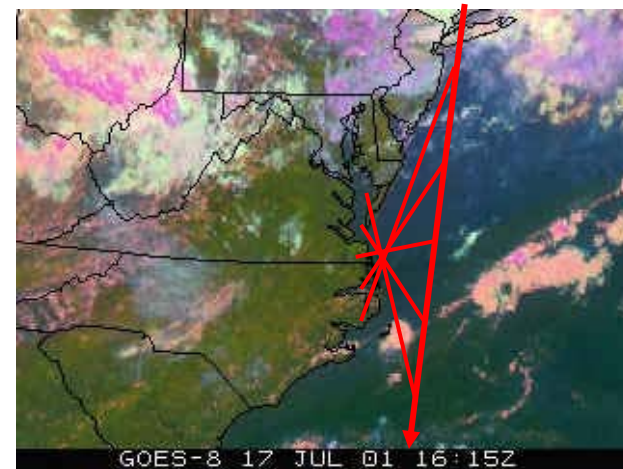
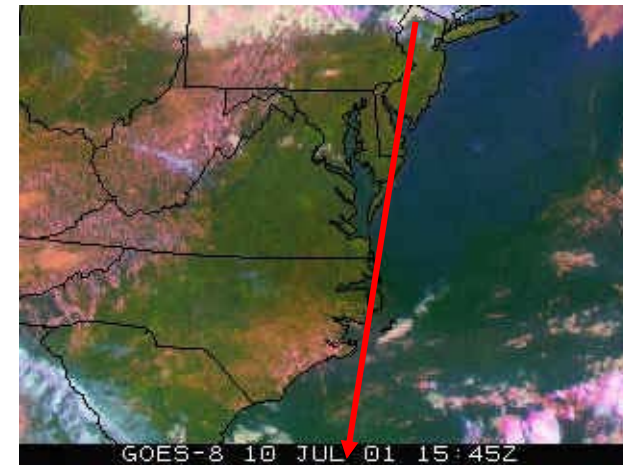
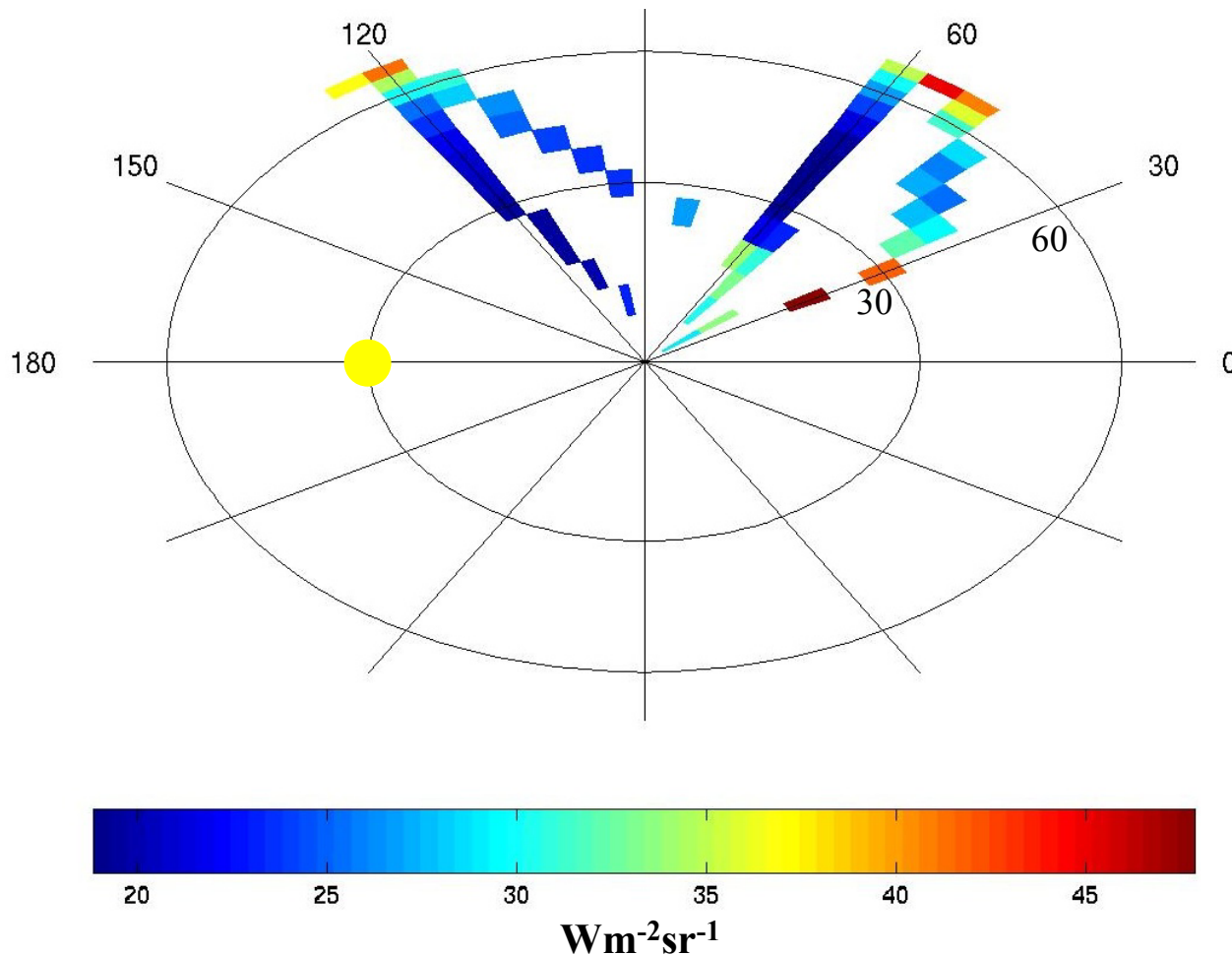
- PAPS + FAPS observations from July 10 to Aug 02, 2001
- Enhanced observation of ocean surface reflectance



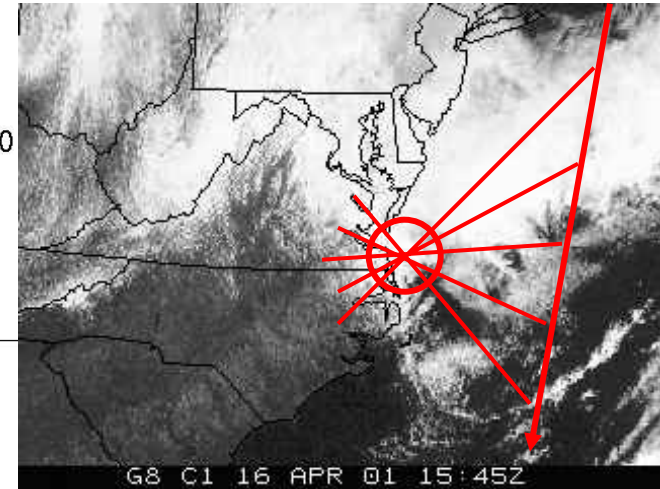
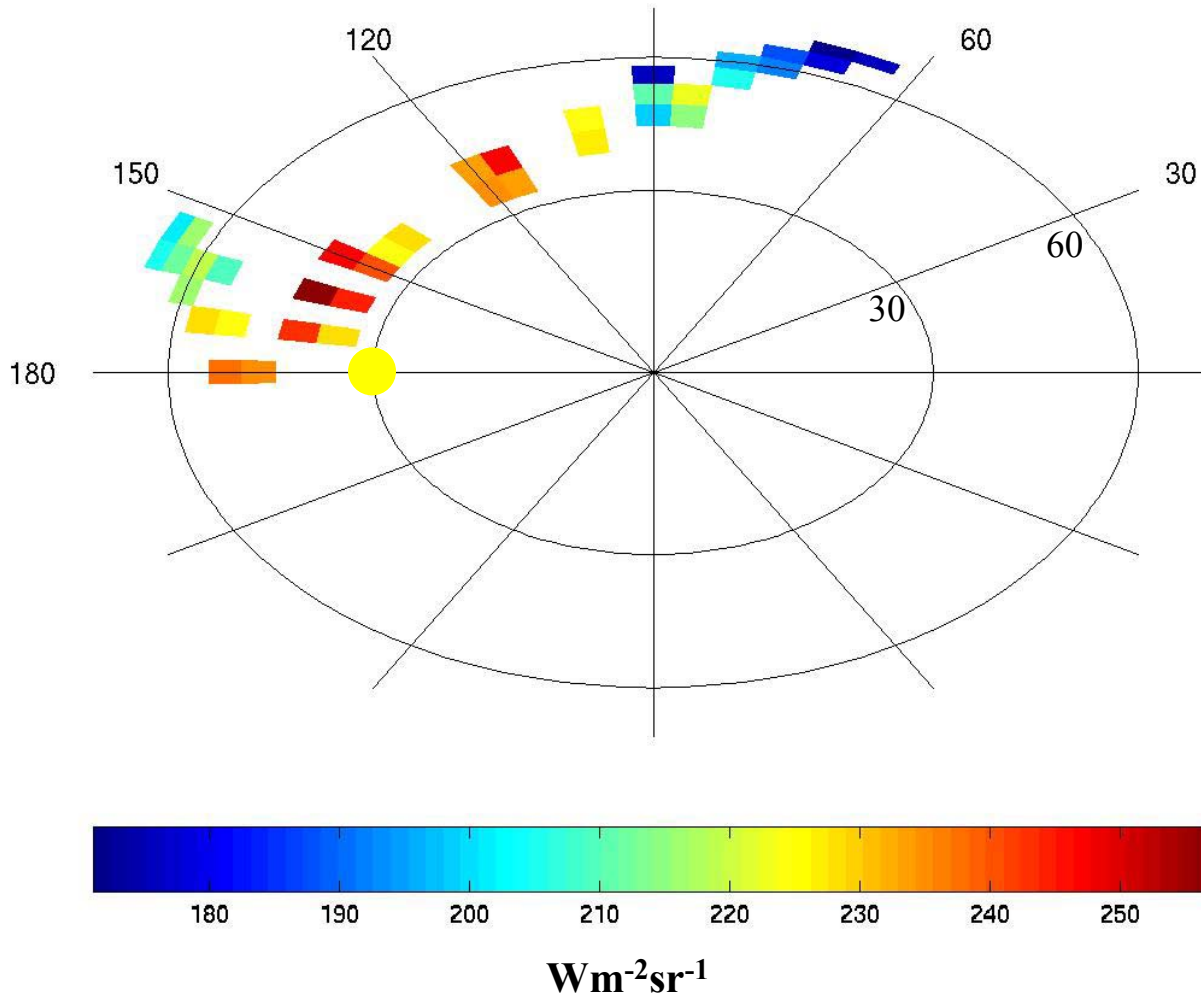
Chesapeake Lighthouse & Aircraft Measurements for Satellites (CLAMS) Jul 10 – Aug 2, 2001



Reflected SW Radiances Clear-Sky Ocean (CLAMS July 10 + 17, 2001)



Reflected SW Radiances Overcast Coastal Region



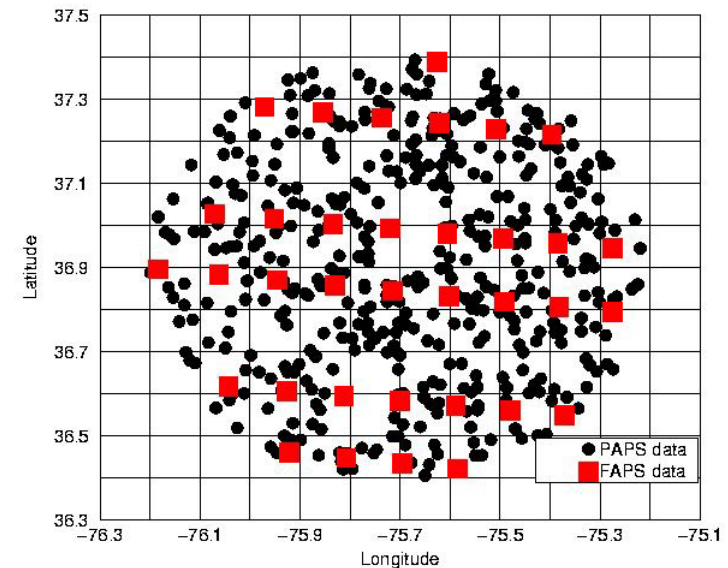
Spatial vs Angular Variability

	Flux (Wm ⁻²)	σ (Wm ⁻²)	D= σ /F (%)	N
FAPS	724.8	46.4	6.4	35
PAPS	676.9	54.8	8.1	447

$$D^2(\text{Angular}) = D^2(\text{Total}) - D^2(\text{Spatial})$$

PAPS ➡ Total + FAPS ➡ Spatial

$$D(\text{Angular}) = 4.9\%$$



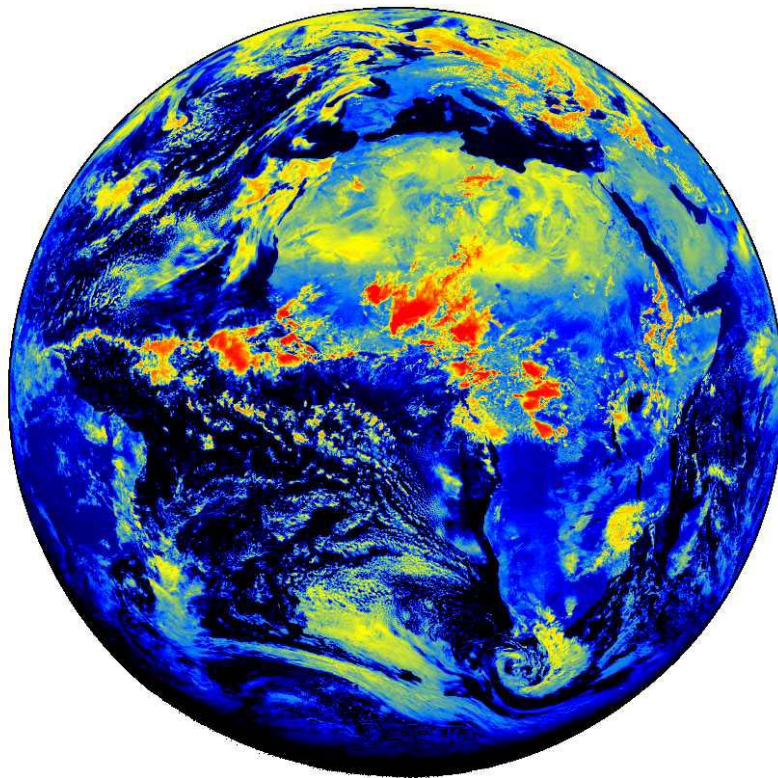
Inter-Calibration of GERB and CERES

GERB:

- 50-km resolution
- 256 detectors



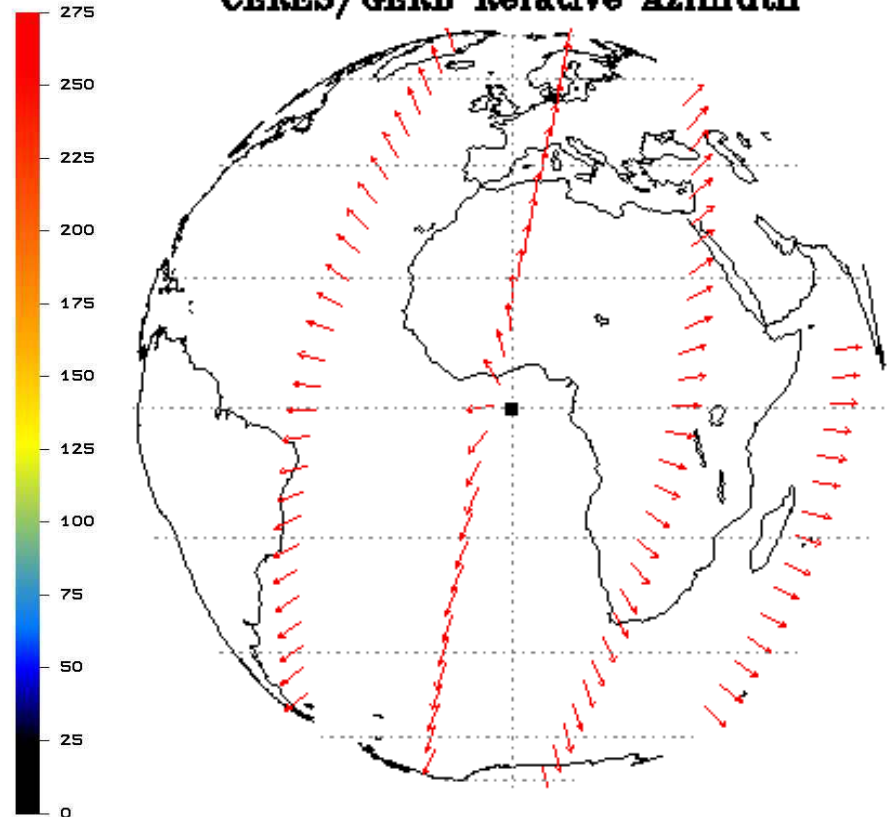
Reflected Solar Radiance ($\text{W/m}^2/\text{sr}$)



CERES:

- 20-km at nadir
- 1 scanning detector

CERES/GERB Relative Azimuth



GERB-like image from Meteosat-7 data by RMIB GERB team

Inter-Calibration of GERB and CERES

- In RAPS mode the azimuth of CERES and GERB are matched 11% of the time ($\pm 10^\circ$ tolerance: $20/180^\circ$)
- In PAPS mode the azimuth of CERES and GERB are matched continuously (100%)
- Spatial noise due to FOV size differences can be reduced by averaging data over 1-deg regions
- In PAPS mode each Terra orbit provides 1 independent comparison of CERES and all GERB detectors

95% confidence interval in the comparison (using 1-deg spatial average)

(%)	Sample/ orbit	30 days		75 days	
		SW	LW	SW	LW
PAPS	120	0.5	0.1	0.3	<0.1
RAPS	13	1.5	0.3	1.0	0.2



Future CERES PAPS Activities

- Intensive sampling of particular relative azimuth planes
- Inter-calibration with GERB
- Multi-angle observation of deep convective clouds during CRYSTAL-FACE
- The CERES team welcomes suggestions for use of the enhanced spatial and angular sampling mode for specific applications
- Visit <http://asd-www.larc.nasa.gov/PAPS>

